

Successful observations of (241) Germania occultation on 2014 April 18

This 184 km diameter asteroid was predicted to occult an 11.8 magnitude star in Gemini, 2UCAC 38779497, on Friday 2014 April 18. The maximum length of disappearance of the star was expected to be 8.5 seconds for someone in the centre of the shadow.

The predicted track of the occultation passed NW to SE across the UK at 20:59 UT. An alert was sent to Asteroids & Remote Planets Section (ARPS) members, and a short article and link were placed on the BAA website. The intention was to engage potential observers, and several enquiries were received as a result.

Four observers in the UK successfully monitored the appulse by video and CCD (Birtwhistle, Haymes, Pratt and Talbot) with Alex Pratt (Leeds) obtaining the only positive UK observation since he was close to the centre line with fair weather. David Pettitt's planned observation was interrupted, and Phil Denyer suffered from cloud.

Four observers (Bulder, Winkel, Rutten, Dangl) posted results from continental Europe via the PLANOCULT list server.

Some e-mail comments

John Talbot: It was fun watching the asteroid creep up on the star and I was convinced it was going to be a positive but it was a close miss here...

Alex Pratt: It's been a long time since my last positive. I agree that the asteroid could be clearly seen approaching the star. I saved a couple of short videos during this time.

Tim Haymes: The asteroid and star merged, but there was no occultation for me. When I observed the star visually the previous day in a 20 cm x250, I found the target star was faint and visible only by averted vision.

Alex Pratt also noted the star looked fainter than the UCAC2 catalogue value used in the prediction. The UCAC4 catalogue gave 12.5V

Roger Pickard (VSS): You can get a pretty accurate V magnitude by using the formula in the *JBAA* paper by Dymock & Miles (*JBAA* 119(3), 2009) as long as you have the CMC-14 data. This you can download using *Guide* and I have a spreadsheet that does the conversion for you. The mag is then 12.87V

Results can be seen on the EURASTER.NET website. The four chords (see Figure 2) indicate the time periods when the star was hidden (occulted). This profile is based on the occultation durations and the geographical position of the observers, from their reports. The asteroid

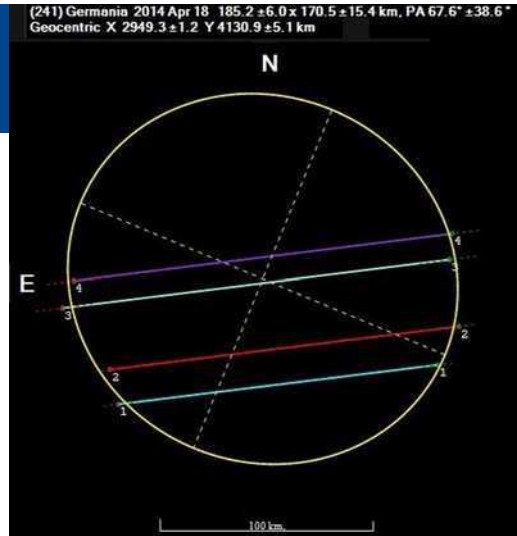


Figure 2. Preliminary chords from four observers outlining the size and shape of (241) Germania.

profile was generated with *Occult 4* software and is preliminary.

Conclusion

The calculated dimensions of the asteroid from these observations are 185±6×171±15 km which is in agreement with previous results of 184 km, and there was a path shift of about 50 km to the NE, with the occultation occurring about 9s earlier than predicted.

Tim Haymes, Asteroid Occultations coordinator, ARPS. [tvh.observatory@btinternet.com]

(5) Astraea to occult an 8th mag star on 2014 August 26

Discovered by K. L. Hencke in 1845, (5) Astraea (123 km diam.) is predicted to occult HIP 4297, an 8.2 magnitude star in Cetus, on Aug 26 between 01:32 and 01:35 UT (02:32–02:35 BST). The expected path crosses Scotland, Northern England, Ireland, Isle of Man and parts of West Wales.

Finder charts and updates are on Steve Preston's web site:

http://www.asteroidoccultation.com/2014_08/0826_5_32908.htm

There is a 0.23 path width uncertainty. In the case of an occultation, the combined light of the asteroid and the star will drop by 3.3 mag to 11.5 mag (the magnitude of the asteroid) for at most 17.7 seconds. The star will be in the south at +35° altitude and should be visible in most instruments, including mounted binoculars under good conditions.

Accounts and impressions welcomed with disappearance and reappearance timings if possible.

Tim Haymes, ARPS
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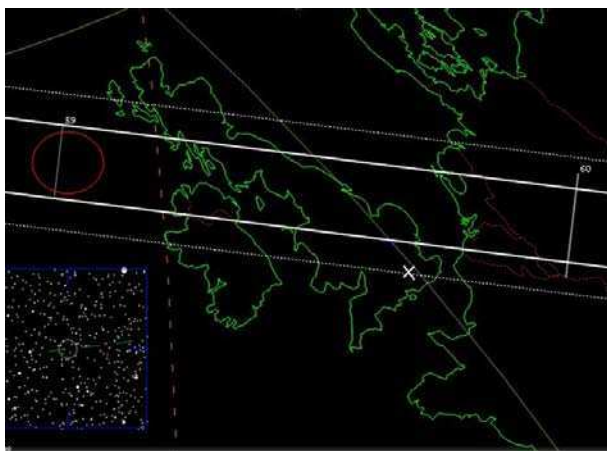


Figure 1. The track of the occultation across the UK plotted with *Occult-4* software.

The prediction, by Steve Preston, could be found at http://www.asteroidoccultation.com/2014_04/0418_241_32426.htm

It was expected that the target star could be seen visually in a 20 cm or larger instrument. This was generally not the case, since it became apparent that the published prediction gave a brighter magnitude than reality. A 20 cm Newtonian was just able to pick up the star by averted vision, and CCD images supported this. The star was nearer 12.5 V (not 11.8). This put the star outside the range of visual detection for most observers. Richard Miles indicated a 25 cm or larger telescope would be needed to monitor the brightness of the star visually with any certainty.

Observers submitting a result to PLANOCULT, or via email

Observer	Location	Result	Method/Timing	Equipment
A. R. Pratt	Leeds	7.04s	Video/GPS	20 cm Mak, WAT-910/HX
J. Talbot	Abingdon	MISS	Video/GPS	30 cm Newt, WAT-120N
T. Haymes	Maidenhead	MISS	Video/GPS	30 cm Newt, WAT-910/HX
P. Birtwhistle	Gt Shefford	MISS	Drift scan/Net	40 cm SCT, U47 CCD
P. Denyer	Hornchurch	Clouded	Video/GPS	23 cm SCT, WAT-120N
H. Bulder	NL	MISS	Eye/DCF77	C12+Stopwatch
J. M. Winkel	Zeddum, NL	8.53s	Video/GPS	31 cm Newt, WAT-120N
H. Rutten	Arcen, NL	7.7s	Video/GPS	35 cm
G. Dangl	Nonndorf, AUT	8.32s	Video/GPS	25 cm Newt, WAT-120N
D. Pettitt	Carlisle	Interrupted	Video	WAT-910HX